TIME ASPECTS OF THE EUROPEAN COMPLEMENT TO GPS: CONTINENTAL AND TRANSATLANTIC EXPERIMENTAL PHASES

P. Uhrich, B. Juompan, R. Tourde Laboratoire primaire du temps et des fréquences [LPTF] Observatoire de Paris, Bureau national de métrologie Paris, France M. Brunet, J–F. Dutrey

> Centre national d'études spatiales [CNES] Toulouse, France

Abstract

The CNES project of a European Complement to GPS [CE-GPS] is conceived to fulfill the needs of Civil Aviation for a non-precise approach phase with GPS as sole navigation means. This generates two missions: a monitoring mission – alarm of failure –, and a navigation mission – generating a GPS-like signal on board the geostationary satellites. The host satellites will be the Inmarsat constellation. The CE-GPS missions lead to some time requirements, mainly the accuracy of GPS time restitution and of monitoring clock synchronization.

To demonstrate that the requirements of the CE-GPS could be achieved, including the time aspects, an experiment has been scheduled over the last two years, using a part of the Inmarsat II F-2 payload and specially designed ground stations based on 10 channels GPS receivers. This paper presents a review of the results obtained during the continental phase of the CE-GPS experiment with two stations in France, along with some experimental results obtained during the transatlantic phase (three stations in France, French Guyana, and South Africa). It describes the synchronization of the monitoring clocks using the GPS Common-view or the C- to L-Band transponder of the Inmarsat satellite, with an estimated accuracy better than 10 ns (1 σ).

INTRODUCTION

The 'Centre national d'études spatiales' (CNES, France) is the French Space Agency. The CNES project of a European Complement to GPS (CE-GPS in the following) is dedicated to the needs of the Civil Aviation community to achieve the requirements of a non-precision approach phase with GPS used as sole navigation means. Many functions have to be fullfilled by such a system, for which the time requirements are reaching the state of the art of the techniques used by the Time Metrology community. The whole CE-GPS project started more than four years ago, and the experimental part more than two years ago.





At this point of the CE–GPS project, all experimental stages have been performed, with a great amount of collected data to be processed. The results presented in this paper concern only the time aspects of the experiment, the calibrations of the ground stations and the synchronization of the monitoring clocks during the continental and transatlantic phases. Both Common-view GPS and Two-way time transfer through geostationary satellite have been used, the processing of the data being carried out by the 'Laboratoire primaire du temps et des fréquences' (LPTF, France). The results obtained are compared to the requirements of the CE–GPS project.

SHORT REMINDER OF THE CE–GPS PROJECT

The concept of the CE-GPS and the experimental system were presented at the EFTF 93^[1]. The CNES project of a European Complement to GPS is mainly dedicated to the needs of the Civil Aviation community. It can be considered as the first step of a French design for a Global Navigation Satellite System (GNSS). The reference mission adopted by the CNES will:

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- achieve the requirements of a non-precision approach phase.
- enable GPS to be used as sole navigation means.

The functions to be fullfiled by such a system are:

- a monitoring mission: alarm of a failure on a GPS satellite within 10 seconds.
- a navigation mission: to increase the GPS availability by eliminating coverage gaps.

Following the proposals of satellite operators, it was agreed that the space segment would be provided by the Inmarsat III geostationary satellites. Because one of the system specifications is to minimize modifications on the existing GPS receivers, the signal transmitted by the geostationary payload shall be similar to a GPS signal. In addition it has to carry specific CE-GPS informations. This and other specifications have many consequences, among which only the time aspects are described in this paper.

To prove the feasability of such a Complement to GPS, and to help estimating the performances and the limits of an operational system, the CNES has organized an experiment of which main objectives are:

- to confirm the ability of transmitting a GPS-like signal from a geostationary satellite.
- to demonstrate the feasability of synchronizing with the GPS time a virtual clock on board the geostationary satellite.
- to demonstrate the capability of GPS receivers to process the CE-GPS signal.
- to evaluate the User Equivalent Range Error (UERE) [2] when using the geostationary satellite.
- to synchronize the ground stations following the requirements.

The space segment of the CE-GPS experiment, beside the GPS NAVSTAR constellation, is the Inmarsat II F2 geostationary satellite of which part of a transponder in the payload was made available free of charge to the CNES for the duration of the experiment. The ground segment is made of three specially designed ground stations based on 10 channels navigation GPS receivers, whose development has been entrusted to the IN-SNEC (Caen, France), and of a computer processing station for the collected data located at the CNES space center of Toulouse (France).

The CE-GPS experimentation was divided into three stages: a first stage in May-June 1993, with two stations located at the IN-SNEC (Caen, France) in parallel with the same clock, called the "calibration phase", a second stage from September to November 1993 with one station at the LPTF (Paris, France) and another one at the CNES space center of Toulouse (France), called the "continental phase"; a third stage from May to June 1994, with three stations located at the CNES space centers of Toulouse (France), Kourou (French Guiana) and Hartebeeshoek (South Africa), called the "transatlantic phase". The acronyms used for these three stations are TLS, KRU, and HBK, with obvious meanings.

TIME ASPECTS OF THE CE–GPS

The time requirements concerning the navigation mission to be fullfilled by the system are the synchronization with respect to GPS satellites. The standard deviation of the time difference between an event issued from the geostationary payload and an equivalent event issued from any GPS satellite should be less or equal to 120 ns with Selective Availability (S.A.) on [2]. With S.A. off, these requirements drop to 20 ns. Concerning the monitoring mission, the time requirements deal with the relative synchronization of the monitoring clocks, which should be within 10 ns (1 s) in accuracy if they are in view of the same geostationary satellite, or within 15 ns (1 s) if not.

The description of the system architecture, of the ground stations, and the discussion about time and frequency servo-control techniques, or orbitography aspects, have been made elsewhere [1,3]. Only the experimental set-up and some of the calibration results are presented here. The method for restituting the GPS time following the requirements, based on the statistical behaviour of the S.A. noise, is described in a paper presented at the EFTF 94 [4], along with some experimental results obtained with a four-channel time dedicated GPS receiver: over an averaging period of 2 h 24 min, and with the simultaneous use of the four channels, it has been demonstrated that the GPS time could be restituted on the ground with an accuracy of 14 ns (1 s). A possible method for achieving a clock synchronization is to use the results of the GPS time restitution separately calculated in remote stations [1,4].

The time aspects of the CE–GPS presented in this paper are the synchronization of the ground stations clocks, either by GPS Common–view or by Two-way satellite time transfer (TWSTT) through the geostationary payload. The well known method for the synchronization of remote atomic clocks is the GPS Common–View technique [5]. Because GPS time dedicated receivers are included in the CE–GPS ground stations, it was decided to use the Common–View technique with the BIPM schedule as the reference for clock synchronization, provided that a calibration of the remote receivers is done, and that atmospheric measurements are made available. Until

now, TWSTT was performed using Eutelsat or Intelsat telecommunication system, or domestic satellites. When using MITREX Modems over short bases (800 km), an accuracy of 1.7 ns (1 s) has been estimated, and directly compared to the equivalent GPS Common-view results [6]. It was proposed to use the Inmarsat C- to L-Band transponder, and the spare C/A GPS gold pn-codes [1,2].

EXPERIMENTAL SET-UP

The description of the ground stations will be limited to the basic equipments involved and to the items and techniques related to the results presented in this paper. The complete presentation of the stations of the CE-GPS experiment can be found elsewhere [1].

Inmarsat II F2 is located -15.5 E. The Sagnac effect for all links is easy to determine with an uncertainty within 0.01 ns (1 s). All stations are identical in terms of a spatial link:

Antenna diameter C-Band: 0.6 m Uplink frequency: 6428.475 MHz (C-Band) Uplink S/C G/T: -14 dB/K Antenna diameter L-Band: 1.2 m Downlink frequency: 1533.475 MHz (L-Band) Downlink on axis G/T: 1.3 dB/K Maximum EIRP: 39.8 dBW

At the start of the experiment, no data were available concerning TWSTT performances when using an Inmarsat transponder, C/A gold codes, and GPS receivers as Modem. The basic sampling period of the data inside these GPS receivers is 0.6 s. It was decided to schedule four sessions per day, each lasting 15 min, to allow statistical analysis on a sufficient amount of data, and to detect any influence of atmospheric parameters on the performances. The sessions took place at 1:15, 7:15, 13:15, and 19:15 TU.

Inside each ground station (figure 1) are implemented two SERCEL NR106, which are tenchannel GPS navigation receivers. They are related, thanks to the switcher 1, either to a common GPS antenna (L1 carrier), or to the receiving antenna of the Inmarsat signal (L-Band) converted to L1 before the switcher. The GPS antenna is also connected to a GPS receiver SERCEL NRT2, which is a 4 channels time dedicated receiver. The NRT2 is directly supplied with the 1pps output of the Cesium clock of the station. This is the classical set-up for the GPS Common-view technique.

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Because the navigation receivers NR106 have no input for dating external events, like the 1 pps from the Cesium clock, it was necessary to build up a so-called 'GPS signals generator', which has two functions. It generates a sequence of C/A code synchronized with the 1 pps output of the clock, which modulates a L1 carrier in order to be dated by the internal counter of both NR106. This signal is denoted '1 pps L1-C/A'. The C/A code chosen for this internal link for all CE-GPS stations is numbered 33: it is a spare for the operational GPS, not to be used until further notice. For refering any external signal to the Cesium clock of the station, the dating of this internal 1 pps L1-C/A signal has to be done simultaneously with the dating of the external signal. The other function of this 'generator' is to output the servo-controlled signal related to

the 1 pps of the clock to the transmitting antenna towards Inmarsat (C-Band carrier). This is similar to the transmitting part of a Modem used for the TWSTT, like MITREX for instance. The equivalent to the receiving part of a classical TWSTT Modem are the NR106 receivers.

There are opportunities for test links, short loop, and changes of the role of each NR106 receiver. There are other possible station configurations to be considered [1], but for the time aspects of the CE-GPS, a stable configuration was chosen for the whole period of data collection. The Cesium clocks monitoring the CE-GPS stations were a HP 5071 A option 1 at the LPTF, a HP 5061 A option 4 at TLS and KRU, and an Oscilloquartz at HBK. All types have proven to remain stable enough to evaluate properly the performances of the CE-GPS stations.

Either the GPS or the C- to L-Band radiowave techniques need estimations of the atmospheric delays. Models for tropospheric delays are working reasonnably well at the nanosecond level. For ionospheric delays, ionospheric calibrators are needed. Ten channels codeless receivers (not on figure 1) were used at all CE-GPS stations for measuring the ionospheric delays on the GPS satellites signals. A polynomial mapping method was scheduled to be used to determine the ionospheric delays in the direction of the Inmarsat geostationary satellite. But many problems occured during the running of the ionospheric calibrators. Considering the small distance between both stations involved in the continental phase (800 km), it was decided to use the STANAG results as the ionospheric delays in the direction of GPS satellites, and a Bent model for the C- and L-Band ionospheric delays in the direction of the Inmarsat satellite. Because the lines of sight of the Inmarsat satellite from both TLS and LPTF stations were very close, the difference of the C-Band ionospheric delays was negligible at the nanosecond level.

For the transatlantic phase, because the ionospheric calibrators have worked well only part of the time, the LPTF proposed to compute the ionospheric delays as following:

- for each 15 s sampled measurements in the direction of GPS satellites, a VTEC (vertical total electronic content) is calculated.
- a mean value of these VTECs is estimated, and projected in the direction of the geostationary satellite.
- for each TWSTT 15 min session, a mean value of the ionospheric delays is computed.
- during the whole periods where TWSTT measurements have been made continuously an average value of the ionospheric delays in the direction of the geostationary satellite is computed, separately for the 4 daily sessions. Three periods of 5 to 7 days duration, where the CE-GPS stations have worked continuously, have been identified during the transatlantic phase of the experiment.

It was proposed to consider as an uncertainty on these values the highest standard deviation of the computed average values, which was 5.1 ns (1 s). This is of course not a state of the art value. It was also proposed to try to use some IGS (International GPS Service for geodynamics) post-processed data, but the results are not available yet.

CALIBRATIONS

All internal delays of the ground stations were either measured (cables) or estimated (electronic components) following the manufacturers data sheets with a good uncertainty. A short loop performed during the calibration stage of the experiment, along with TWSTT sessions with both stations connected to the same clock, have shown that the measured differential delays were in good agreement with the estimated values, given the estimated global uncertainty. One of the critical issues is the calibration of the 'GPS signal generators'. Examining the synchronization equations, it appears that the 1 pps signal group delay through each generator must be monitored. Frequent calibration sessions were scheduled in all stations during the whole experiment. A calibration consists roughly in measuring the time delay between the arrivals into the NRT2 of two homologous 1 pps signals: a direct 1 pps signal from the Cesium clock and the 1 pps L1–C/A signal supplied through the generator output. This procedure was possible because the NRT2 have proven to remain quite stable all over the experiment.

The results of the successive calibrations during the continental phase are presented on figure 2. The calibrations at CNES station appear more stable than the LPTF station measurements. Moreover the CNES station was switched off for a while on MJD 49266 (October 6): there are obviously two sets of data before and after this MJD, each scattered within 4 ns. On the contrary the LPTF data are much more scattered within 25 ns. It has been shown by the manufacturer of the stations that, owing to the components used for this experiment, the signal generator could exhibit stepwise varying delays, each step being a multiple of 2.44 ns. By processing in deferred time the dating of the 1 pps L1–C/A performed continuously by both NR106 of each station, those variations could be identified, summed and compared to the day to day calibrations of figure 2. With the analysis of these records, some improvements could be achieved, but a lot of discrepencies are remaining, mainly due to missing data. This is why these discontinuities in the delays are disregarded. The average values of the calibration sessions are adopted along with the relevant standard deviations as uncertainties, giving for the signal generators delays:

- at TLS (continental phase): mean value = 1007.5 ns; standard deviation = 4.8 ns.

- at LPTF: mean value = 998.5 ns; standard deviation = 7.1 ns.

A similar behaviour of the signal generators was observed during the transatlantic phase of the experiment (Figure 3). The generator of the HBK station, formerly located at the LPTF, remained the most perturbated one, compared to the generator of the KRU station, which has never been switched off during the whole experiment. The manufacturer of the stations believes that the generator of the KRU station has reached the best achievable stability, with the components used for the building of it. The average values of the calibration sessions were adopted along with the relevant standard deviations as uncertainties, as for the continental phase, giving for the signal generators delays:

- at TLS (transatlantic phase): mean value = 1009.4 ns; standard deviation = 6.3 ns.

- at HBK: mean value = 986.1 ns; standard deviation = 7.6 ns.
- at KRU: mean value = 1025.8 ns; standard deviation = 1.4 ns.

A calibration by transportation of two GPS receivers was organized during the continental phase of the experiment, so that three round-trips were performed between the LPTF and the TLS station. It has worked remarquably well, except for some local problems too long to explain here, giving the usual uncertainty of 1.5 ns (1 s) on the differences between the two time scales. An other control of the results was possible thanks to the operational GPS Common-view link between the two stations involved, used continuously for the computation of the TAF ('Temps atomique francais').

No temperature effect was noticeable on the sets of data.

CONTINENTAL EXPERIMENTAL PHASE

An analysis of the data by the Modified Allan deviation showed that white phase noise is preponderant on each 15 min session over a period long enough to allow the computation of an average value. The precision of one 0.6 s sampled CE-GPS observation is about 9 to 10 ns, as can be deduced from the Modified Allan deviation. It could be compared to the MITREX observations that are ten times better [6], but it remains consistent with the magnitude expected when using C/A code along with standard discriminators on signals affected by a measured C/No of about 40 dB(Hz).

On Figure 4 are presented the average values of the TWSTT 15 min sessions, compared to the Common-view GPS daily averages between the GPS time receivers connected to the CE-GPS stations. One can see that the TWSTT points are well distributed around the GPS curve, even though they are scattered. It appears that many sessions have suffered from the above mentionned stepwise varying delays compared to the average values of the generators delays. This is a logical consequence of the choice to use an average of the calibration measurements. The uncertainties of the generators delays are obviously the most degrading part for the accuracy of this experiment of TWSTT through Inmarsat. We propose the following uncertainty budget in nanoseconds (1 s):

Inmarsat transponder	0	[Global coverage]
Sagnac effect	0.0	
Ionospheric delays	2.0	[Differential/Bent model]
Differential delays (LPTF-CNES)	6.2	[1ppsL1–C/A generator]
UTC(LPTF)-1 pps REF	0.5	[short line]
UTC(CNES)-1 pps REF	1.0	[long line]
Two-way accuracy	6.6	[Quadratic sum]

This accuracy is given for each session averaged result, and remain well below the upper limit of the CE-GPS requirements (10 ns). It surely will drop to a more interesting value compared to the MITREX results [6] if the behaviour of the generators could be withdrawn from the measurements. The uncertainty of the comparison between GPS Common-view and TWSTT can be estimated in nanoseconds (1 s) by:

TWSTT (CE–GPS continental)	6.6	
GPS Common-view [5]	3.4	[800 km baseline]
Uncertainty of the comparison	7.5	

This uncertainty appears consistent with the data plotted on Figure 4.

TRANSATLANTIC EXPERIMENTAL PHASE

A statistical analysis of the transatlantic data showed that a white phase noise behaviour is observed, as for the continental phase. Until now, no GPS Common-view data are available to be compared to the TWSTT results, and a lot of TWSTT data are missing, due to local problems in the equipments. On figures 5, 6 and 7 are plotted the TWSTT results on each link. The regular behaviour of the two different types of atomic clocks used can yet be recognized. In a very similar way as for the continental phase, we propose the following uncertainty budgets in nanoseconds (1 s) for each link:

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Ionospheric delays	5.1 [average	computed from	GPS measurements]
Equipment delays	TLS - HBK	HBK - KRU	KRU - TLS
7.0	5.5	4.6	
UTC(k)-1 pps REF		1.0	[two stations]
TWSTT accuracies	8.7	7.6	6.9

These accuracies are near the upper limit of the CE-GPS requirements (10 ns). One of the possible tests of the consistency of the results, along with a check of the computation options, is to calculate the deviation from the closure between the three stations. For this purpose, daily independent TWSTT results have been built up, for two of the three TWSTT links, by interpolation between two separated measurements sessions, centered on the date of the third link result. The deviations from the closure are plotted on figure 8. Even if only few points are available, there is obviously a bias between the three different sets of data. Again the behaviour of the generators of the stations is suspected to be the most important part of this bias. The results are scattered between -8.0 ns and 15.3 ns. These values are consistent with the estimated accuracies given above. But the weak number of computed values do not allow us to realize a more complete statistical analysis.

If one could consider that the behaviour of the 1 pps L1–C/A generator in the KRU station would be the regular behaviour of an operational CE–GPS ground station, and if the ionospheric delays could be measured with a better uncertainty, by using the IGS post–processed data for instance, one could estimate the following uncertainty budget:

Ionospheric delays	2.0
Equipment delays	1.5
UTC(k)-1 pps REF	1.0 [two stations]
TWSTT accuracy	2.7

This could be the best achievable accuracy of the method, with similar stations more than 10,000 km apart, as far as the global coverage of the satellite's antenna is realized.

CONCLUSION

It has been demonstrated that the time requirements of the CE–GPS could be fulfilled with ground stations comparable to those built up for the purpose of the experiment presented in this paper, either for a distance between the stations inferior to 1000 km, or for stations more than 10,000 km apart. Even with the calibration problems related to the signal generators as conceived for this experiment, the monitoring clock synchronization could be done by the TWSTT technique through an Inmarsat geostationary satellite, with C/A gold codes and GPS receivers, with an accuracy of 7 to 9 ns (1 s). The best achievable accuracy is estimated within 2.7 ns (1 s).

The software for computing the timing data from the pseudo-range measurements made by the NR106, which are only GPS navigation receivers, is nearly completed. It will allow in the future the computation of GPS Common-view data the same way as in GPS receivers dedicated to time measurements, making the NRT2 receivers connected to the CE-GPS stations useless. Because NR106 are 10 channels receivers, it will supply data for the restitution of GPS time with the highest number of GPS satellites available simultaneously, reducing the averaging period for a similarly reduced S.A. noise.

Beside TWSTT, other techniques are scheduled to be tested with signals transmitted by the Inmarsat transponder. Among others, the servo-control could be the most interesting due to its near real-time time transfer capability. In this case however the synchronization is less accurate than the best achievable with post-processed data.

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Figure 1. Diagram of a CE-GPS ground station.



Figure 2. Calibration of the generators at the LPTF (*) and at TLS (*) during the continental phase of the experiment.



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Figure 3. Calibration of the generators at the KRU (°), the TLS (+), and the HBK (Δ) stations during the transatlantic phase of the experiment.



Figure 4. TWSTT 15 min averaged sessions and GPS common-view daily averages (CE-GPS continental phase).



Figure 6. TWSTT 15 min averaged sessions between TLS and HBK (CE-GPS transatlantic phase).