# Heterodyne frequency measurements on the 11.6- $\mu$ m band of OCS: new frequency/wavelength calibration tables for 11.6- and 5.8- $\mu$ m OCS bands

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Heterodyne difference frequency measurements between a  $^{13}\text{CO}_2$  laser and a diode laser tuned (and in most cases locked) to the peaks of OCS absorption lines have been used to improve frequency calibration tables in the  $860\text{-cm}^{-1}$  region by factors of 20-50. Measurements have been made on the vibrational transitions  $10^{00}\text{-}00^{00}$ ,  $11^{10}\text{-}01^{10}$ , and  $20^{00}\text{-}10^{00}$  for OCS. The measurements on the  $10^{00}\text{-}00^{00}$  and  $20^{00}\text{-}10^{00}$  transitions are also used to provide frequency calibration tables for the  $20^{00}\text{-}00^{00}$  band of OCS near  $1700\text{ cm}^{-1}$ .

## I. Introduction

In an earlier paper heterodyne measurements were reported for some of the frequency differences between the  $2\nu_2$  absorption band of OCS (carbonyl sulfide) near  $1050~\rm cm^{-1}$  and the nearby  $\rm CO_2$  laser transitions. Since the frequencies of the  $\rm CO_2$  laser transitions have been measured with high accuracy, such heterodyne measurements give very accurate frequencies for the OCS absorption lines. Even more recent measurements by Sattler et~al. on higher rotational transitions give further improvements in the accuracy of the OCS absorption lines near  $1050~\rm cm^{-1}$ . Such measurements are needed to provide convenient and accurate absorption frequency standards for the calibration of tunable laser devices, such as tunable diode lasers.

This paper reports the results of similar difference frequency measurements between  $^{13}\text{CO}_2$  laser lines and the  $\nu_1$  OCS absorption lines near 870 cm $^{-1}$ . Intensity measurements also have been made for two  $\nu_1$  OCS absorption lines. From these measurements a table of intensities and recommended frequencies (with the estimated uncertainties) is given for use in the calibration of the frequencies of tunable laser devices.

An earlier paper<sup>4</sup> reported the results of grating and diode laser measurements on the  $\nu_1$  and  $2\nu_1$  bands of OCS. That paper gave preliminary wave numbers for OCS lines from 825 to 885 cm<sup>-1</sup> and from 1655 to 1737 cm<sup>-1</sup>. While those tables are still useful for indicating the relative positions and intensities of some of the hot bands and isotopic species, this paper presents tables of line frequencies (and wave numbers) that are more than an order of magnitude more accurate.

# II. Description of Experimental Details

Since a detailed account of our heterodyne measurement technique was presented earlier, we will give only a recapitulation and indicate some recent changes. A block diagram of the experimental arrangement for our frequency measurement procedure is shown in Fig. 1. The tunable diode laser (TDL) is operated in a liquid helium Dewar which can accommodate three such devices. The chopped tunable diode laser beam goes through a monochromator, and a portion of the emerging beam is split off and goes through an OCS cell to a detector. The signal from the detector is phase detected at the chopper frequency and displayed on a recorder. In addition to the obvious functions of mode rejection and coarse calibration, the monochromator is also used to level the transmitted power as a function of frequency as the TDL scans in the vicinity of the OCS line to be measured. A flat background level allows one to use a first derivative (instead of a third derivative lock) to stabilize the diode laser to the OCS line.

This work also gives heterodyne measurements and conventional diode laser measurements for a number of hot band absorption lines. Measurements on the  $20^{0}0-10^{0}0$  hot band yield constants which have been used to provide improved calibration tables for the  $20^{0}0-00^{0}0$  band of OCS near  $1710 \text{ cm}^{-1}$ .

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Received 4 December 1980.

To continue on with the frequency measurement, the chopper is removed from the system, and the diode laser is frequency modulated at  $\sim$ 1 kHz. The phase detected signal (first derivative) is fed back to the laser control module in the proper phase to lock the diode laser frequency to the OCS line. For transitions up to P(46), the  $^{13}\text{CO}_2$  laser was stabilized by locking to a saturated absorption dip in low pressure (5.3-Pa or 40-mTorr) isotopic  $\text{CO}_2$  following the scheme of Freed and Javan. For J > 46, an unstabilized 2-m laser was carefully tuned to line center. Our frequency measurements of P(52) and P(54) of this laser (which contained the same

laser fill pressure for all our experiments) agreed well with values (predicted for stabilized lasers) which are used in Table I.

The HgCdTe mixer used to combine these two laser signals had a 3-dB bandwidth slightly greater than 100 MHz. The heterodyne beat notes generally did not have adequate SNR for cycle counting; hence a marker from an oscillator was superposed on the heterodyne beat note and simultaneously measured on a frequency counter.

The potential accuracy of the frequency measurements depends on several factors. One factor is how

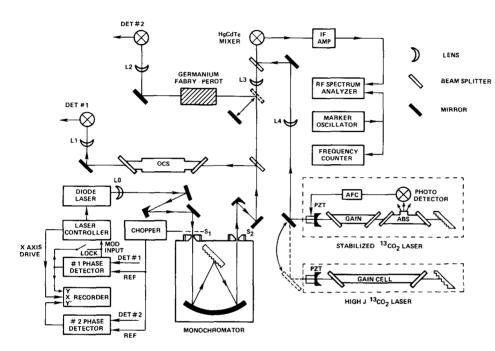


Fig. 1. Block diagram of a TDL spectrometer with capabilities for heterodyne frequency measurements with <sup>13</sup>CO<sub>2</sub> lasers. A kinematically mounted mirror permits the user to select the germanium etalon for wavelength metrology or the heterodyne configuration for frequency metrology. A second mirror mount arrangement permits selection of the appropriate <sup>13</sup>CO<sub>2</sub> laser.

Table I. Frequencies of 11.5-μm Band Carbonyl Sulfide Absorption Lines Measured by the Heterodyne Technique

Vibrational transition	Rot. trans.	OCS- <sup>13</sup> CO <sub>2</sub> meas. freq. diff. (MHz)	Measured OCS freq. (MHz) <sup>a</sup>	Obscalc. (MHz)	$ m CO_2$ laser $^b$ transition	$^{13}\mathrm{CO}_2\mathrm{freq.^c}$
1000-000	R(87)	+73.0	26 674 291.6 (6.0)	0.4	<sup>13</sup> P(28)	26 674 218.62
$10^{0}0-00^{0}0$	R(75)	+2773.8	26 566 226.9 (3.0)	-0.1	P(32)	26 563 453.07
$10^{0}0-00^{0}0$	R(57)	+395.6	26 393 239.6 (2.0)	-0.1	P(38)	26 392 843.97
$10^{0}0-00^{0}0$	R(40)	+1362.7	26 218 193.2 (3.0)	0.0	P(44)	26 216 830.49
$10^{0}0-00^{0}0$	R(18)	+1733.6	25 975 344.4 (2.0)	-1.8	P(52)	25973610.78
$10^{0}0-00^{0}0$	R(7)	-1668.0	25 847 177.1 (2.0)	1.5	P(25)	25 848 845.09d
$10^{0}0-00^{0}0$	P(1)	-1277.5	25 739 018.1 (2.0)	0.8	P(29)	$25\ 740\ 295.57^d$
11 <sup>1e</sup> 0-01 <sup>1e</sup> 0	R(84)	+174.0	26 450 485.0 (3.0)	0.3	P(36)	26 450 311.04
$11^{1e}0-01^{1e}0$	R(42)	-628.4	26 034 711.5 (6.0)	0.2	P(50)	26 035 339.88
$11^{1e}0-01^{1e}0$	R(25)	-707.3	25 848 137.8 (2.0)	-0.7	P(25)	25 848 845.09d
$11^{1e}0-01^{1e}0$	R(22)	-1400.2	25 814 164.2 (2.0)	1.1	P(26)	$25\ 815\ 564.36^{d}$
$11^{1f}0-01^{1f}0$	R(42)	+820.7	26 036 160.6 (6.0)	-1.0	P(50)	26 035 339.88
$11^{1f}0-01^{1f}0$	R(25)	-34.5	25 848 810.6 (4.0)	-0.2	P(25)	$25\ 848\ 845.09^{d}$
$11^{16}0-01^{16}0$	R(22)	-841.4	25 814 723.0 (4.0)	-1.9	P(26)	25 815 564.36 <sup>d</sup>
$20^{0}0-10^{0}0$	R(86)	+600.4	26 450 911.4 (3.0)	0.1	P(36)	26 450 311.04
$20^{0}0-10^{0}0$	R(73)	-439.6	26 334 336.9 (3.0)	-0.3	P(40)	26 334 776.55
$20^{0}0-10^{0}0$	R(50)	+810.7	26 111 131.5 (4.0)	1.2	P(15)	$26\ 110\ 320.77^d$
$20^{0}0-10^{0}0$	R(40)	+85.0	26 007 511.5 (4.0)	0.0	P(19)	$26\ 007\ 426.54^{d}$
$20^{0}0-10^{0}0$	R(31)	-329.0	25 910 930.4 (6.0)	2.3	P(54)	25 911 259.42
$20^{0}0 - 10^{0}0$	R(16)	+2772.1	25 743 067.7 (3.0)	-0.7	P(29)	$25\ 740\ 295.58^d$

<sup>&</sup>lt;sup>a</sup> The estimated uncertainty in megahertz is given in parentheses.

<sup>&</sup>lt;sup>b</sup> A 1.25-m stabilized <sup>13</sup>CO<sub>2</sub> laser was used for lines up to P(46). For J above P(46) and for hot band transitions, a 2-m laser was used.

<sup>&</sup>lt;sup>c</sup> Except for hot band lines, laser reference frequencies were taken from Freed et al.<sup>7</sup>

<sup>&</sup>lt;sup>d</sup> Recent NBS measurements of lasing <sup>13</sup>CO<sub>2</sub> hot band transitions (01<sup>1</sup>1-11<sup>1</sup>0<sub>I</sub>).

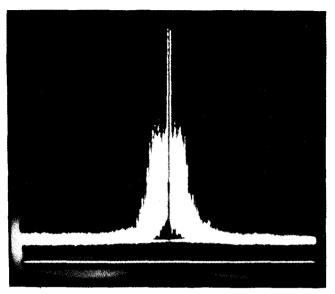


Fig. 2. Heterodyne beat note between P(52) of the  $^{13}\mathrm{CO}_2$  laser and the TDL when locked to R(30) of the  $O^{13}\mathrm{CS}$  molecule. The center frequency is 516.9 MHz, and the sharp feature is due to the marker oscillator. The 10-cm long spectrum analyzer trace has a 10-MHz/cm dispersion.

accurately one can determine the OCS line center. We have taken the uncertainty in finding the line center to be equal to the full width at half-maximum (FWHM) divided by the ratio of the derivative SNR. At 11  $\mu$ m, the OCS FWHM is ~40 MHz. Hence a SNR of 50 will permit one to locate the line center with an uncertainty less than 1 MHz. The frequency modulation producing the derivative signal also broadens the nominal 3-MHz signal up to ~10 MHz, as shown in Fig. 2. We consider this a near ideal beat note and believe that its center can be located within 1 MHz. The narrow feature in the center is due to the marker oscillator.

Figure 3 shows the overlap of the <sup>13</sup>CO<sub>2</sub> laser and the 11.6-um OCS absorption band. The vertical lines indicate some close coincidences for potential heterodyne measurements. The solid part of the <sup>13</sup>CO<sub>2</sub> laser curve represents the operable range of our 1.25-m stabilized CO<sub>2</sub> laser. When the <sup>13</sup>CO<sub>2</sub> laser was stabilized, its frequency uncertainty was <0.2 MHz, and thus this frequency uncertainty was small compared with other possible contributions to uncertainty in the OCS frequency measurements shown in Table I. The dashed portion of the curve represents the extended range with the unstabilized 2-m higher gain laser. When the 2-m laser is used, we estimate that with great care one can tune its frequency to within ~1 or 2 MHz of the line center. For the key measurements involving the P(29)hot band laser transition, the 2-m laser was stabilized with an external absorption cell.

The OCS absorption cell had a path length of 50 cm and pressures ranged from 12 Pa (0.09 Torr) for the stronger 10<sup>0</sup>0–00<sup>0</sup>0 lines to 266 Pa (2 Torr) for some of the 20<sup>0</sup>0–10<sup>0</sup>0 lines. For the weaker 20<sup>0</sup>0–10<sup>0</sup>0 lines, a 2-m absorption cell heated to 155°C was needed.

Several unexpected developments occurred during these experiments. First, the measurement of the OCS

R(18) line with respect to the P(52) <sup>13</sup>CO<sub>2</sub> laser line indicated that the frequency of P(52) in the literature<sup>6</sup> was off considerably. New work by Freed and coworkers has verified this frequency discrepancy. Our <sup>13</sup>CO<sub>2</sub> frequency values are taken from their new paper.<sup>7</sup> Second, our attempt to reach P(60) with the 2-m laser by modifying it with a new discharge tube and high reflectivity mirror did not succeed. P(56) was the limit in the new laser, and this excluded the potential measurement of R(2) with respect to P(58). As a result of these improvements, however, we did observe some eighteen lasing hot band  $(01^{1}1-11^{1}0_{\text{I}})$  transitions in this region ranging from P(10) to P(30) and have measured their frequencies relative to the stabilized CO<sub>2</sub> laser.<sup>2</sup> These hot band measurements were made using a metal-insulator-metal or MIM diode and will be described later. 8 One hot band transition, P(29), was used to obtain the only heterodyne measurement in the P branch of the OCS band.

Only a small number of lines of the OCS hot bands could be measured using the heterodyne technique because of the limited frequency response of our mixer. To cover a wider range of rotational quantum numbers and thereby improve the constants for the levels involved in the hot band transitions, we have also measured the absorption spectrum of OCS using a tunable diode laser and etalon.

For these nonheterodyne measurements, a closed cycle cooler was used, and more extensive temperature tuning of the diode was possible. A solid germanium etalon with a free spectral range of 0.0162 cm<sup>-1</sup> was used to measure the wave number differences between the hot band lines and the calibration lines of the 10<sup>0</sup>0–00<sup>0</sup>0 band. The new values presented at the end of this paper were used for the calibration of these measurements. Reference 4 more clearly describes how these measurements were made. These (OCS) hot band measurements are given in Table II.

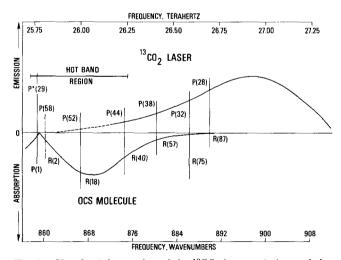


Fig. 3. Sketch of the overlap of the  $^{13}\text{CO}_2$  laser emission and the 11.6- $\mu$ m OCS absorption bands as a function of frequency/wave number. Only a qualitative indication of the intensities is shown in order to better illustrate the extent of the bands.

### III. Results and Analysis of the Data

Table I gives the results of the present heterodyne measurements. In addition to the measured frequency differences and the derived frequencies of the OCS lines, Table I also gives the frequencies for the  $^{13}\mathrm{CO}_2$  laser transitions determined by the most recent work.  $^{7,8}$ 

The assignment of the measured OCS lines was fairly straightforward. The mode selecting monochromator was calibrated so that the central frequency of a given diode scan was known to within about  $\pm 0.10\,\mathrm{cm^{-1}}$ . The tables given in Ref. 4 and also calculations provided by Fayt<sup>9</sup> were extremely useful for assigning lines through pattern recognition, i.e., by taking into account frequency separations and relative intensities. When needed for line identification, solid germanium etalons with fringe spacings (free spectral ranges) of 0.017 or 0.05 cm<sup>-1</sup> were used to measure approximate frequency differences.

When it was suspected that an absorption line was due to one of the isotopes (<sup>18</sup>O<sup>12</sup>C<sup>32</sup>S, <sup>16</sup>O<sup>13</sup>C<sup>32</sup>S, or <sup>16</sup>O<sup>12</sup>C<sup>34</sup>S), the spectrum of an enriched sample was

recorded to verify the assignment. In this way a few measured lines were shown to be due to <sup>16</sup>O<sup>13</sup>C<sup>32</sup>S, although the vibrational assignment could not always be determined (further measurements on isotopic species will be the subject of a future paper).

To analyze the two l=0 vibrational transitions,  $10^{0}0-00^{0}0$  and  $20^{0}0-10^{0}0$ , we used the term value expression:

$$T(v,J) = F(v) + B_v J(J+1) - D_v J^2 (J+1)^2 + H_v J^3 (J+1)^3,$$
 (1)

$$\nu_0(v' \leftarrow v'') = F(v') - F(v''). \tag{2}$$

In addition to the heterodyne measurements reported in Table I, the analysis included appropriate microwave measurements taken from Refs. 10 and 11. The analysis of the  $20^{0}0-10^{0}0$  band also included diode measurements calibrated against the  $10^{0}0-00^{0}0$  lines (using Table III) and given in Table II. A least squares analysis of the data was performed with each datum given a weight inversely proportional to the square of the estimated uncertainty. The uncertainties assigned to each measurement are given in Tables I and II.

Table II. Hot Band Transitions for OCS Calibrated Against the 10°0-00°0 Frequencies

	$11^{1e}0-01^{1e}0$			$11^{16}0-01^{16}0$			$20^{0}0-10^{0}0$			
Rotational transition	Obs. freq. (unc.) <sup>a</sup> (MHz)	O-C <sup>b</sup> (MHz)	Rotational transition	Obs. freq. (unc.) <sup>a</sup> (MHz)	O-C <sup>b</sup> (MHz)	Rotational transition	Obs. freq. (unc.) <sup>a</sup> (MHz)	O-C <sup>b</sup> (MHz		
P(41)	25 016 835 (18)	5	P(41)	25 017 087 (18)	-10	P(41)	25 015 363 (12)	-15		
P(33)	25 123 957 (24)	2	P(33)	25 124 041 (24)	-7	P(33)	25 123 049 (32)	-8		
P(31)	25 150 399 (34)	-8	P(31)	25 150 435 (34)	-29	P(31)	25 149 580 (41)	-34		
P(14)	25 369 904 (12)	8	P(8)	25 444 996 (72)	-1	P(15)	25 356 839 (12)	6		
P(8)	25 445 056 (72)	-15	P(7)	25 457 428 (50)	13	P(14)	25 369 499 (30)	24		
P(7)	25 457 473 (50)	-10	$P(5)^c$	25 482 167 (56)	13	P(8)	25 444 543 (68)	-10		
$P(5)^{e}$	25 482 221 (56)	13	$R(1)^c$	25 567 748 (50)	32	P(7)	25 456 940 (50)	3		
$R(1)^c$	25 567 720 (50)	33	R(4)	25 603 886 (20)	-9	P(6)	25 469 294 (58)	9		
R(4)	25 603 835 (20)	19	R(17)	25 757 254 (50)	12	P(5)	25 481 612 (52)	16		
R(17)	25 756 864 (50)	18	R(18)	25 768 802 (40)	-4	R(1)	25 566 772 (20)	30		
R(18)	25 768 376 (40)	-2	R(19)	25 780 332 (40)	-4	R(3)	25 590 728 (12)	-8		
R(19)	25 779 882 (40)	6	R(20)	25 791 832 (36)	-1	R(17)	25 754 538 (72)	14		
R(20)	25 791 337 (32)	-3	R(21)	25 803 299 (32)	3	R(18)	25 765 927 (20)	-15		
R(21)	25 802 774 (28)	5	R(24)	25 837 502 (20)	19	R(19)	25 777 313 (16)	-10		
R(27)	25 870 620 (12)	5	R(27)	25 871 367 (12)	1	R(20)	25 788 678 (12)	13		
R(28)	25 881 796 (12)	<b>-</b> 5	R(37)	25 982 092 (40)	0	R(21)	25 799 968 (12)	-2		
R(37)	25 980 890 (48)	-7	R(43)	26 046 877 (64)	6	R(22)	25 811 240 (20)	3		
R(45)	26 066 567 (12)	-4	R(44)	26 057 532 (64)	-13	R(24)	25 833 656 (12)	-1		
R(46)	26 077 132 (20)	13	R(45)	26 068 174 (20)	-10	R(28)	25 878 044 (60)	3		
R(47)	26 087 634 (20)	3	R(46)	26 078 790 (16)	$^2$	R(38)	25 986 322 (12)	3		
R(62)	26 240 939 (72)	9	R(62)	26 243 601 (50)	10	R(44)	26 049 428 (40)	-1		
R(63)	26 250 847 (12)	<del>-</del> 7	R(63)	26 253 584 (28)	1	R(45)	26 059 810 (48)	-1		
R(64)	26 260 731 (20)	-10	R(64)	26 263 537 (20)	-2	R(46)	26 070 147 (36)	-6		
R(65)	26 270 567 (20)	-24	R(65)	26 273 460 (20)	3	R(47)	26 080 469 (28)	12		
R(66)	26 280 400 (35)	-3	R(66)	26 283 345 (35)	7	R(63)	26 239 922 (80)	7		
R(67)	26 290 177 (12)	0	R(68)	26 303 002 (12)	12	R(65)	26 259 124 (30)	-2		
R(68)	26 299 914 (20)	1	R(69)	26 312 748 (12)	-11	R(66)	26 268 673 (16)	3		
R(69)	26 309 618 (22)	6		. ,		R(67)	26 278 161 (50)	-13		
•						R(68)	26 287 652 (32)	16		
						R(69)	26 297 051 (40)	-8		
						R(70)	26 306 443 (48)	3		

<sup>&</sup>lt;sup>a</sup> The estimated uncertainty (primarily associated with temperature instability in the etalon which we believe is the major source of error in these experiments) is given in parentheses in megahertz.

<sup>&</sup>lt;sup>b</sup> Observed frequencies minus frequencies calculated with the constants given in Table III.

<sup>&</sup>lt;sup>e</sup> The separation of the e and f components for P(5) and R(1) was estimated, and the unresolved measurements were corrected for this splitting.

Table III. Wave Numbers, Frequencies, and Intensities (at 296 K) of Spectral Lines of the 1000-0000 Band of OCS from 815 to 892 cm<sup>-1</sup>

			Intensity $(m^{-2} daPa^{-1})$				Intensity (m <sup>-2</sup> daPa <sup>-1</sup>
Wave number (cm <sup>-1</sup> )	Frequency (MHz)	Tran- sition	$(cm^{-2} atm^{-1})^a$	Wave number (cm <sup>-1</sup> )	Frequency (MHz)	Tran- sition	or (cm <sup>-2</sup> atm <sup>-1</sup> )
815.09994(40)	24436081.4(119)	P(95)	0.00031	848.02784(4)	25423235.1(12)	P(26)	0.34621
815.61764(36)	24451601.6(107)	P(94)	0.00037	848.46352(4)	25436296.4(12)	P(25)	0.35059
816.13413(32)	24467085.8(96)	P(93)	0.00044	848.89801(4)	25449322.0(12)	P(24)	0.35377
816.64943(29)	24482534.0(86)	P(92)	0.00052	849.33130(4)	25462311.9(12)	P(23)	0.35565
817.16353(26)	24497946.3(77)	P(91)	0.00061	849.76341(4)	25475266.0(12)	P(22)	0.35616
817.67643(23)	24513322.7(68)	P(90)	0.00073	850.19432(4)	25488184.5(12)	P(21)	0.35524
818.18813(20)	24528663.1(61)	P(89)	0.00086	850.62404(4)	25501067.2(12)	P(20)	
818.69864(18)	24543967.7(54)	P(88)	0.00101	851.05257(4)	25513914.1(12)	P(19)	
	24559236.3(48)	P(87)	0.00101	851.47990(4)	25526725.2(12)	P(18)	
819.20794(16)		P(86)	0.00119	851.90604(4)	25539500.5(12)	P(17)	
819.71606(14)	24574469.2(42)						
820.22297(12)	24589666.2(37)	P(85)	0.00163	852.33098(4)	25552240.0(13)	P(16)	
820.72870(11)	24604827.3(33)	P(84)	0.00191	852.75473(4)	25564943.6(13)	P(15)	
821.23322(10)	24619952.7(30)	P(83)	0.00223	853.17728(4)	25577611.4(13)	P(14)	0.30476
821.73656(9)	24635042.3(27)	P(82)	0.00259	853.59863(4)	25590243.2(13)	P(13)	0.29106
822,23870(8)	24650096.1(24)	P(81)	0.00301	854.01879(4)	25602839.2(13)	P(12)	0.27579
822.73965(7)	24665114.1(22)	P(80)	0.00349	854.43775(4)	25615399.2(13)	P(11)	
823.23940(7)	24680096.4(21)	P(79)	0.00404	854.85550(4)	25627923.3(13)	P(10)	
823.73797(7)	24695043.0(20)	P(78)	0.00466	855.27206(4)	25640411.4(13)	P(9)	0.22110
824.23534(7)	24709953.9(20)	P(77)	0.00536	855.68742(4)	25652863.4(13)	P(8)	0.20016
824.73152(6)	24724829.0(19)	P(76)	0.00616	856.10157(4)	25665279.4(13)	P(7)	0.17801
325.22651(6)	24739668.5(19)	P(75)	0.00707	856.51452(4)	25677659.4(13)	P(6)	0.15478
825.72032(6)	24754472.3(19)	P(74)	0.00809	856.92627(4)	25690003.3(13)	P(5)	0.13059
826.21293(6)	24769240.5(19)	P(73)	0.00924	857.33681(4)	25702311.0(13)	P(4)	0.10556
826.70435(6)	24783973.0(19)	P(72)	0.01052	857.74615(4)	25714582.6(13)	P(3)	0.07984
327.19459(6)	24798669.8(19)	P(71)	0.01196	858.15428(4)	25726818.0(13)	P(2)	0.05357
827.68363(6)	24813331.1(19)	P(70)	0.01357	858.56120(4)	25739017.3(13)	P(1)	0.02690
828.17149(6)	24827956.7(19)	P(69)	0.01536	859.37142(4)	25763306.9(13)	R(0)	0.02698
828.65816(6)	24842546.7(19)	P(68)	0.01735	859.77471(4)	25775397.4(13)	R(1)	0.05389
829.14364(6)	24857101.1(19)	P(67)	0.01956	860.17679(4)	.25787451.4(13)	R(2)	0.08055
	24871619.9(19)	P(66)	0.02200	860.57766(4)	25799469.1(13)	R(3)	0.10682
829.62794(6)			0.02260	* *			
830.11105(6)	24886103.2(19)	P(65)		860.97731(4)	25811450.5(13)	R(4)	0.13255
830.59297(6)	24900550.9(18)	P(64)	0.02764	861.37575(4)	25823395.3(13)	R(5)	0.15757
831.07371(6)	24914963.0(18)	P(63)	0.03088	861.77297(4)	25835303.7(13)	R(6)	0.18176
831,55326(6)	24929339.5(17)	P(62)	0.03443	862.16898(4)	25847175.6(13)	R(7)	0.20498
832.03162(6)	24943680.6(17)	P(61)	0.03829	862.56376(4)	25859011.0(13)	R(8)	0.22710
832,50880(6)	24957986.0(17)	P(60)	0.04250	862.95733(4)	25870809.8(13)	R(9)	0.24802
832.98480(5)	24972256.0(16)	P(59)	0.04706	863.34967(4)	25882571.9(13)	R(10)	0.26762
833.45961(5)	24986490.4(16)	P(58)	0.05199	863.74079(4)	25894297.4(13)	R(11)	0.28582
833.93323(5)	25000689.3(15)	P(57)	0.05731	864.13069(4)	25905986.2(13)	R(12)	0.30255
834.40567(5)	25014852.7(15)	P(56)	0.06304	864.51936(4)	25917638.3(13)	R(13)	0.31773
334.87692(5)	25028980.5(14)	P(55)	0.06917	864.90680(4)	25929253.6(13)	R(14)	0.33131
835,34699(5)	25043072.9(14)	P(54)	0.07573	865.29302(4)	25940832.1(12)	R(15)	0.34326
835.81588(4)	25057129.7(13)	P(53)	0.08272	865.67801(4)	25952373.7(12)	R(16)	0.35355
836.28358(4)	25071151.0(13)	P(52)	0.09014	866.06176(4)	25963878.4(12)	R(17)	0.36218
336.75010(4)	25085136.8(13)	P(51)	0.09800	866.44428(4)	25975346.2(12)	R(18)	
837.21543(4)	25099087.1(12)	P(50)	0.10630	866.82557(4)	25986777.0(12)	R(19)	
337.67958(4)	25113001.9(12)	P(49)	0.11503	867.20563(4)	25998170.7(12)	R(20)	0.37815
338.14254(4)	25126881.2(12)	P(48)	0.12417	867.58445(4)	26009527.4(12)	R(21)	0.38027
338.60432(4)	25140725.0(12)	P(47)	0.13372	867.96203(4)	26020846.9(12)	R(22)	0.38085
339.06491(4)	25154533.3(11)	P(46)	0.14366	868.33837(4)	26032129.3(12)	R(23)	0.37997
339,52432(4)	25168306.1(11)	P(45)	0.15396	868.71346(4)	26043374.5(12)	R(24)	0.37768
339.98255(4)	25182043.3(11)	P(44)	0.16459	869.08732(4)	26054582.3(12)	R(24) $R(25)$	0.37407
340.43959(4)	25195745.1(11)	P(43)	0.17552	869,45993(4)	26065752.9(12)	R(26)	
340.89545(4)	25209411.3(11)	P(42)	0.18671	869.83129(4)	26076886.1(12)	R(20) R(27)	0.36321
	25223042.0(11)	P(42) = P(41)	0.19811	870.20141(4)	26087981.9(12)	R(27)	0.35613
341.35012(4)			0.19811 $0.20966$		26087981.9(12) 26099040.3(12)		
341.80360(4)	25236637.2(11)	P(40)		870.57027(4)		R(29)	0.34810
842.25591(4)	25250196.8(11)	P(39)	0.22131	870.93789(4)	26110061.1(12)	R(30)	0.33920
342.70702(4)	25263720.9(11)	P(38)	0.23300	871.30425(4)	26121044.3(12)	R(31)	0.32953
343.15695(4)	25277209.5(11)	P(37)	0.24465	871.66936(4)	26131990.0(12)	R(32)	0.31919
343,60569(4)	25290662.5(11)	P(36)	0.25619	872.03321(4)	26142897.9(12)	R(33)	0.30829
344.05325(4)	25304079.9(12)	P(35)	0.26754	872.39580(4)	26153768.2(11)	R(34)	0.29692
344.49962(4)	25317461.8(12)	P(34)	0.27862	872.75713(4)	26164600.6(11)	R(35)	0.28517
344.94481(4)	25330808.1(12)	P(33)	0.28933	873.11720(4)	26175395.2(11)	R(36)	0.27313
845.38881(4)	25344118.8(12)	P(32)	0.29959	873.47601(4)	26186152.0(11)	R(37)	0.26090
845.83162(4)	25357393.9(12)	P(31)	0.30930	873.83355(4)	26196870.8(11)	R(38)	0.24855
846.27324(4)	25370633.4(12)	P(30)	0.31836	874.18982(4)	26207551.6(11)	R(39)	0.23617
346.71367(4)	25383837.3(12)	P(29)	0.32668	874.54483(4)	26218194.3(11)	R(40)	0.22382
	25397005.6(12)	P(28)	0.33416		,	/	
847.15292(4)	20001000.0(12)	1 (20)	0.00110				

Table III, continued

		114	Intensity
			(m <sup>-2</sup> daPa <sup>-1</sup> )
Wave number	Frequency	Tran-	or
(cm <sup>-1</sup> )	(MHz)	sition	$(cm^{-2} atm^{-1})^a$
074 9095 674)	00000709 0(11)	D(41)	0.01157
874.89856(4)	26228798.9(11)	R(41)	0.21157
875.25102(4)	26239365.4(11)	R(42)	0.19949
875.60220(4)	26249893.6(11)	R(43)	0.18762
875.95211(4)	26260383.5(11)	R(44)	0.17602
876.30073(4)	26270835.0(12)	R(45)	0.16474
876.64808(4)	26281248.2(12)	R(46)	0.15380
876.99414(4)	26291622.9(12)	R(47)	0.14324
877.33892(4)	26301959.0(12)	R(48)	0.13308
877.68240(4)	26312256.5(13)	R(49)	0.12335
878.02460(4)	26322515.4(13)	R(50)	0.11407
878.36551(4)	26332735.6(13)	R(51)	0.10523
878.70512(5)	26342916.9(14)	R(52)	0.09685
879.04344(5)	26353059.4(14)	R(53)	0.08893
879.38046(5)	26363163.0(15)	R(54)	0.08147
879.71618(5)	26373227.6(15)	R(55)	0.07447
880.05060(5)	26383253.2(16)	R(56)	0.06791
880.38371(5)	26393239.6(16)	R(57)	0.06179
880.71551(6)	26403186.8(17)	R(58)	0.05609
881.04601(6)	26413094.8(17)	R(59)	0.05081
881.37519(6)	26422963.5(17)	R(60)	0.04591
881.70306(6)	26432792.8(18)	R(61)	0.04140
882.02961(6)	26442582.6(18)	R(62)	0.03725
882.35485(6)	26452332.9(19)	R(63)	0.03344
882.67876(6)	26462043.6(19)	R(64)	0.02995
883.00135(6)	26471714.6(19)	R(65)	0.02677
883.32262(6)	26481345.9(19)	R(66)	0.02388
883.64255(6)	26490937.3(19)	R(67)	0.02125
883.96116(6)	26500488.9(19)	R(68)	0.01887
884.27843(6)	26510000.5(19)	R(69)	0.01671
884.59437(6)	26519472.1(19)	R(70)	0.01478
884.90897(6)	26528903.6(19)	R(71)	0.01304
885.22223(6)	26538294.9(19)	R(72)	0.01148
885.53415(6)	26547646.0(19)	R(73)	0.01008
885.84472(6)	26556956.7(19)	R(74)	0.00884
886.15395(7)	26566227.0(20)	R(75)	0.00773
886.46182(7)	26575456.8(20)	R(76)	0.00674
886.76834(7)	26584646.1(21)	R(77)	0.00587
887.07351(7)	26593794.8(22)	R(78)	0.00511
887.37732(8)	26602902.7(24)	R(79)	0.00443
887.67976(9)	26611969.8(27)	R(80)	0.00383
887.98085(10)	26620996.1(29)	R(81)	0.00331
888.28057(11)	26629981.5(33)	R(82)	0.00285
888.57892(12)	26638925.8(37)	R(83)	0.00245
888.87590(14)	26647829.0(42)	R(84)	0.00210
889.17150(16)	26656691.0(48)	R(85)	0.00180
889.46573(18)	26665511.8(54)	R(86)	0.00154
889.75858(20)	26674291.2(61)	R(87)	0.00131
890.05005(23)	26683029.2(68)	R(88)	0.00112
890.34013(26)	26691725.7(77)	R(89)	0.00095
890.62883(29)	26700380.7(86)	R(90)	0.00080
890.91614(32)	26708993.9(96)	R(91)	0.00068
891.20205(36)	26717565.5(107)	R(92)	0.00057
891.48657(40)	26726095.1(119)	R(93)	0.00048
891.76970(44)	26734582.9(131)	R(94)	0.00041
892.05142(48)	26743028.7(145)	R(95)	0.00034

 $<sup>^{\</sup>alpha}$  To within the accuracy of the intensity data, 1 m $^{-2}\,daPa^{-1}$  is equal to 1 cm $^{-2}\,atm^{-1}.$ 

Since the data span the range of rotational quantum numbers from J=0 to J=88, the H term was tentatively included in Eq. (1). In agreement with earlier conclusions,  $^{10}$  it was found that the ground state data could be fit quite well with  $H_0=0$ . On the other hand, the data were fit somewhat better when H was included as a variable for the fits involving the  $20^{00}$  and  $10^{00}$  states. Even though the uncertainty in the H term is very large, it was included in the analysis of transitions involving those states.

Since one goal of this work was to obtain data for the calculation of the  $20^{0}0-00^{0}0$  transitions near 1700 cm<sup>-1</sup>, the  $10^{0}0-00^{0}0$  and  $20^{0}0-10^{0}0$  transitions were analyzed together. This analysis yielded the variance–covariance matrix elements needed to calculate the uncertainties for the calculated  $20^{0}0-00^{0}0$  transitions.

The constants given in Table IV and the variance—covariance matrix elements were used to calculate the frequency calibration data given as Tables III and V. In these tables, the calculated uncertainty in each transition frequency is given in parentheses following each entry. This uncertainty is twice the estimated standard deviation obtained from the statistical analysis. This extra factor of 2 was used somewhat arbitrarily to give what we believe is a better estimate of the absolute uncertainty since the statistically determined uncertainty is always less than the absolute uncertainty in the frequency because of systematic errors.

One possible source of systematic error could be due to a pressure shifting of the OCS absorption lines. The magnitude of the pressure shift is probably about a tenth of the pressure broadening effect which is <75 kHz/Pa (10 MHz/Torr). Since the measurements on the 10<sup>0</sup>0-00<sup>0</sup>0 band were made with OCS pressures smaller than 66 Pa (0.5 Torr), the pressure shift is expected to be <0.5 MHz. Since the  $20^{0}0-10^{0}0$  band measurements were made at pressures as high as 266 Pa (2 Torr), frequency shifts as large as 2 MHz might affect those measurements. However, since the weaker high J transitions are the ones most likely to have smaller pressure shifts, we expect that the errors due to pressure shifts will be smaller than the standard deviations given by the random errors in the measurements. Unfortunately, we are unable to quantify more precisely the sources of systematic error which might affect our results.

The reader is cautioned not to use the constants given in Table IV to calculate transitions beyond the J=95 cutoff that was used for making Tables III and V. Such extrapolations beyond the range of measured transitions are unreliable.

In the analysis of the data for the l=1 vibrational transition (11<sup>1</sup>0–01<sup>1</sup>0), the following term value equations were used:

$$T(v,J) = F(v) + B_v J(J+1) \pm 0.5 q_v J(J+1) - D_v [J(J+1) - 1]^2 \mp 0.5 q_{vJ} J^2 (J+1)^2,$$
 (3)

$$\nu_0(v' \leftarrow v'') = F(v') - F(v''). \tag{4}$$

Table IV. Constants a for OCS Determined from These Measurements

Constant	This work (MHz) <sup>b</sup>	Previous work (Ref. 4) (MHz)
$\nu_0 (20^00-00^00)$	51 293 779.3 (19)	51 293 779. (111)
$\nu_0 (10^00-00^00)$	25 751 180.2 (6)	25 751 204. (65)
$\nu_0 (20^00-10^00)$	25 542 599.0 (18)	<u> </u>
$\nu_0 (11^10 – 01^10)$	25 543 432.7 (14)	25 543 403. (75)
$B(00^00)$	6081.492518 (56)	(6081.4924439)c
D(0000)	$1.3019369 (114) \times 10^{-3}$	$(1.301789 \times 10^{-3})^c$
$B(10^00)$	6063.358265 (200)	6063.3588 (8)
$D(10^{0})$	$1.329873(197) \times 10^{-3}$	$1.3312(19) \times 10^{-3}$
$H(10^{0}0)$	$0.0001214 (276) \times 10^{-6}$	<del></del>
$B(20^{0}0)$	6044.874923 (2550)	6044.891 (5)
$D(20^{0}0)$	$1.359731 (810) \times 10^{-3}$	$1.3607(10) \times 10^{-3}$
$H(20^00)$	$0.0002234 (706) \times 10^{-6}$	
$B(01^{1}0)$	6092.078222 (289)	
$D(01^10)$	$1.323567 (625) \times 10^{-3}$	_
$q(01^10)$	6.361411394 (745)	
$q_J (01^10)$	$4.25608 (667) \times 10^{-6}$	_
$B(11^{1}0)$	6075.50528 (186)	6075.519 (7)
$D(11^{1}0)$	$1.362242 (6q (11^{1}0)$	6.853553 (2121)
6.853 (13)	• • •	,
$q_{J}$ (11 <sup>1</sup> 0)	$11.670 (489) \times 10^{-6}$	

<sup>&</sup>lt;sup>a</sup> The statistically estimated uncertainties in the last digits (one standard deviation) are given in parentheses following each value. Since systematic errors have not been taken into account, these estimated uncertainties may be slightly optimistic.

For both the  $11^{10}$  and  $01^{10}$  states, the upper sign in Eq. (3) applies for the f rotational levels and the lower sign applies for the e rotational levels. As before, the data given in Tables I and II were combined with appropriate microwave transitions  $^{10,11}$  in a weighted least squares fit. Insufficient data were available to warrant extending Eq. (3) to include the higher order H term.

### IV. Description and Analysis of the Intensity Data

Preliminary intensity measurements have been made on several lines of the  $10^{0}0-00^{0}0$  band using the technique of measuring the intensity at the line center, as was done on the  $2\nu_2$  band<sup>1</sup> and as advocated by Jennings.<sup>12</sup> Contrary to our initial expectation, such intensity measurements are not entirely free of instrumental resolution problems.

It was found that some of the measurements implied intensities that were clearly too low. When heterodyne measurements of the frequency width of the diode laser output were possible, it was found that the low intensity measurements were due to large laser linewidths. Since the Doppler halfwidth of the OCS lines is only ~20 MHz, accurate intensity measurements of the absorption peak are only possible if the laser linewidth is <3 MHz. Should the laser linewidth be >3 MHz, the measured peak intensity would be smaller than the true intensity. Our limited experience has indicated that most diode lasers will give unacceptably large output linewidths at certain frequencies although the same diode may have very narrow linewidths at other

frequencies. The best way to ensure that this instrumental problem is not affecting the measurements is to measure simultaneously the diode laser linewidth by beating it against a narrow line gas laser, such as a  $\rm CO_2$  laser.

The most reliable intensity measurements were made on the R(28) and R(33) lines of the  $10^00-00^00$  band. Measurements were made using a 50-cm absorption cell and pressure of the order of 12 Pa (0.09 Torr). The signal representing 100% absorption was determined by blocking the laser beam with an opaque shutter. The absence of stray radiation was verified by observing that the opaque shutter gave the same signal level as was given by saturated absorption lines when the cell was filled to a pressure of 50 Torr of OCS.

At 296 K, the line intensity for R(28) was found to be  $\sim 3.4~\rm cm^{-2}~MPa^{-1}~(0.34~\rm cm^{-2}~atm^{-1})$  and for R(33) it was  $\sim 3.1~\rm cm^{-2}~MPa^{-1}~(0.31~\rm cm^{-2}~atm^{-1})$  with an uncertainty of  $\sim 10\%$ . This gives a transition moment of 0.0632 D and an integrated intensity of  $\sim 350~\rm cm^{-2}~MPa^{-1}~(35~\rm cm^{-2}~atm^{-1})$  for all isotopes and hot bands in the  $\nu_1$  region. This value is close to  $340~\rm cm^{-2}~MPa^{-1}~(34~\rm cm^{-2}~atm^{-1})$  given by Robinson,  $^{13}$  and the transition moment of 0.0632 D is close to the value of 0.0681 D given by Foord and Whiffen.  $^{14}~\rm Tables~III~and~V~give$  estimated line intensities in units of cm $^{-2}~\rm atm^{-1}$  calculated for a temperature of 296 K. To convert the latter intensity units to the frequently used spectroscopic units of cm/molecule, one should divide by 2.479  $\times~10^{19}~\rm molecules \cdot cm^{-3} \cdot atm^{-1}$ .

<sup>&</sup>lt;sup>b</sup> To convert from frequency units to wave number units, use 299 792 458 m/sec for the speed of light. <sup>16</sup>

<sup>&</sup>lt;sup>c</sup> See Ref. 10; the negligible difference in the  $B_0$  and  $D_0$  values given in Ref. 10 and in this work arises from the inclusion of an  $H_0$  term in Ref. 10.

Table V. Wave Numbers, Frequencies, and Intensities (at 296 K) of Spectral Lines of the 2000-0000 Band of OCS from 1662 to 1738 cm<sup>-1</sup>

Table V. Wave				pectral Lines of the 20 C			
Wave number	Frequency	Tran-	Intensity (m <sup>-2</sup> daPa <sup>-1</sup> ) or	Wave number	Frequency	Tran-	Intensity (m <sup>-</sup> 2 daPa <sup>-1</sup> ) or
(cm <sup>-1</sup> )	(MHz)	sition	$(cm^{-2} atm^{-1})^a$	$(cm^{-1})$	(MHz)	sition	$(cm^{-2} atm^{-1})^a$
1001 F000F(40)	40011015 0(140)	D(05)	0.00005	1000 (2007(2)	50052007 4(94)	D(9g)	0.05912
1661.52665(49) 1662.16342(43)	49811315.8(148) 49830405.6(130)	$P(95) \\ P(94)$	0.00005 0.00006	1699.63607(8) 1700.10263(8)	50953807.4(24) 50967794.7(25)	$P(26) \\ P(25)$	0.05912
1662.79765(38)	49849419.5(114)	P(93)	0.00008	1700.56677(8)	50981709.1(25)	P(24)	0.06038
1663.42936(33)	49868357.6(99)	P(92)	0.00009	1701.02847(9)	50995550.7(26)	P(23)	0.06068
1664.05854(29)	49887220.0(86)	P(91)	0.00011	1701.48775(9)	51009319.4(26)	P(22)	0.06075
1664.68520(25)	49906006.8(75)	P(90)	0.00013	1701.94459(9)	51023015.2(27)	P(21)	0.06058
1665.30934(22)	49924718.1(66)	P(89)	0.00015	1702.39900(9)	51036638.2(28)	P(20)	0.06015
1665.93096(19)	49943353.9(57)	P(88)	0.00018	1702.85099(9)	51050188.3(28)	P(19)	0.05946
1666.55007(17)	49961914.3(51)	P(87)	0.00021	1703.30054(10)	51063665.5(29)	P(18)	0.05849
1667.16667(15)	49980399.3(46)	P(86)	0.00024	1703.74766(10)	51077069.8(30)	P(17)	0.05725
1667.78075(14)	49998809.1(42)	P(85)	0.00028	1704.19235(10)	51090401.3(31)	P(16)	0.05574
1668.39233(13)	50017143.7(40)	P(84)	0.00033	1704.63460(11)	51103659.8(32)	P(15)	0.05394
1669.00140(13)	50035403.2(38)	P(83)	0.00039	1705.07443(11)	51116845.4(32)	P(14)	0.05187
1669.60797(12)	50053587.6(37)	P(82)	0.00045	1705.51182(11)	51129958.1(33)	$P(13) \\ P(12)$	0.04953
1670.21203(12) 1670.81360(12)	50071697.1(37) 50089731.6(37)	$P(81) \\ P(80)$	$0.00052 \\ 0.00061$	1705.94678(11) 1706.37931(12)	51142997.9(34) 51155964.7(35)	P(12) = P(11)	$0.04692 \\ 0.04405$
1671.41267(12)	50107691.3(37)	P(79)	0.00070	1706.80940(12)	51168858.6(35)	P(10)	0.04403
1672.00925(12)	50107031.3(37)	P(78)	0.00070	1700.30340(12)	51181679.4(36)	P(9)	0.03758
1672.60323(12)	50143386.3(37)	P(77)	0.00093	1707.66228(12)	51194427.3(36)	P(8)	0.03401
1673.19492(12)	50143380.3(87)	P(76)	0.00107	1708.08507(12)	51207102.1(37)	P(7)	0.03024
1673.78403(12)	50178782.7(37)	P(75)	0.00122	1708.50542(12)	51219703.9(37)	P(6)	0.02629
1674.37064(12)	50196369.1(37)	P(74)	0.00140	1708.92333(13)	51232232.6(38)	P(5)	0.02217
1674.95478(12)	50213881.0(37)	P(73)	0.00160	1709.33881(13)	51244688.2(38)	P(4)	0.01792
1675.53643(12)	50231318.5(37)	P(72)	0.00182	1709.75184(13)	51257070.7(38)	P(3)	0.01355
1676.11560(12)	50248681.6(37)	P(71)	0.00207	1710.16244(13)	51269380.1(38)	P(2)	0.00909
1676.69230(12)	50265970.5(36)	P(70)	0.00235	1710.57059(13)	51281616.3(39)	P(1)	0.00456
1677.26651(12)	50283185.1(36)	P(69)	0.00266	1711.37958(13)	51305869.0(38)	R(0)	0.00457
1677.83826(12)	50300325.5(35)	P(68)	0.00300	1711.78040(13)	51317885.5(38)	R(1)	0.00913
1678.40753(12)	50317391.8(35)	P(67)	0.00338	1712.17879(13)	51329828.7(38)	R(2)	0.01365
1678.97432(11)	50334384.0(34)	P(66)	$0.00380 \\ 0.00426$	1712.57472(13) 1712.96821(12)	51341698.5(38) 51353495.0(37)	R(3) $R(4)$	$0.01810 \\ 0.02245$
1679.53865(11) 1680.10052(11)	50351302.1(34) 50368146.3(33)	$P(65) \\ P(64)$	0.00426	1713.35925(12)	51365218.0(37)	R(5)	0.02243
1680.65991(11)	50384916.6(33)	P(63)	0.00533	1713.74784(12)	51376867.6(36)	R(6)	0.02008
1681.21684(11)	50401613.0(33)	P(62)	0.00594	1714.13397(12)	51388443.7(36)	R(7)	0.03469
1681.77131(11)	50418235.6(32)	P(61)	0.00660	1714.51766(12)	51399946.3(35)	R(8)	0.03842
1682.32332(11)	50434784.4(32)	P(60)	0.00733	1714.89889(12)	51411375.3(35)	R(9)	0.04195
1682.87287(11)	50451259.5(32)	P(59)	0.00811	1715.27766(11)	51422730.6(34)	R(10)	0.04525
1683.41996(11)	50467660.9(32)	P(58)	0.00896	1715.65398(11)	51434012.3(33)	R(11)	0.04831
1683.96460(11)	50483988.6(32)	P(57)	0.00987	1716.02784(11)	51445220.3(32)	R(12)	0.05113
1684.50678(11)	50500242.7(32)	P(56)	0.01085	1716.39923(11)	51456354.5(32)	R(13)	0.05368
1685.04650(11)	50516423.3(32)	P(55)	0.01191	1716.76817(10)	51467414.9(31)	R(14)	0.05596
1685.58378(11)	50532530.4(32)	P(54)	0.01303	1717.13464(10)	51478401.4(30)	R(15)	0.05796
1686.11860(11)	50548564.0(32)	P(53)	0.01423	1717.49864(10)	51489314.0(29) 51500152.7(28)	R(16)	0.05969
1686.65097(10)	50564524.1(31)	$P(52) \\ P(51)$	$0.01550 \\ 0.01685$	1717.86018(9) 1718.21925(9)	51510917.2(28)	$R(17) \\ R(18)$	0.06113 $0.06228$
1687.18090(10)	50580410.9(31)			1718.57585(9)		R(19)	0.06223
1687.70838(10) 1688.23341(10)	50596224.2(31) 50611964.3(31)	P(50) = P(49)	$0.01827 \\ 0.01977$	1718.92997(9)	51521607.7(27) 51532224.1(26)	R(20)	0.06377
1688.75599(10)	50627631.0(31)	P(48)	0.02133	1719.28162(9)	51542766.2(26)	R(21)	0.06411
1689.27614(10)	50643224.5(31)	P(47)	0.02297	1719.63079(8)	51553234.1(25)	R(22)	0.06419
1689.79384(10)	50658744.8(30)	P(46)	0.02467	1719.97748(8)	51563627.6(25)	R(23)	0.06403
1690.30909(10)	50674191.8(30)	P(45)	0.02643	1720.32169(8)	51573946.7(24)	R(24)	0.06363
1690.82191(10)	50689565.7(30)	P(44)	0.02825	1720.66341(8)	51584191.4(24)	R(25)	0.06300
1691.33229(10)	50704866.4(29)	P(43)	0.03011	1721.00265(8)	51594361.5(24)	R(26)	0.06217
1691.84023(10)	50720094.0(29)	P(42)	0.03202	1721.33940(8)	51604457.0(24)	R(27)	0.06114
1692.34573(10)	50735248.5(29)	P(41)	0.03397	1721.67366(8)	51614477.9(24)	R(28)	0.05993 $0.05857$
1692.84879(9)	50750329.9(28)	P(40)	0.03594	1722.00543(8)	51624424.0(24) 51634295.3(24)	$R(29) \\ R(30)$	0.05857 0.05705
1693.34941(9)	50765338.3(28) 50780273.6(27)	P(39) = P(38)	$0.03793 \\ 0.03992$	1722.33470(8) 1722.66147(8)	51634295.3(24) 51644091.8(25)	R(30) $R(31)$	0.05541
1693.84760(9) 1694.34336(9)	50780273.6(27) 50795136.0(27)	P(38) = P(37)	0.03992	1722.98575(8)	51653813.2(25)	R(31)	0.05341
1694.83668(9)	50809925.3(26)	P(36)	0.04130	1723.30752(9)	51663459.7(26)	R(33)	0.05182
1695.32756(9)	50824641.7(26)	P(35)	0.04580	1723.62678(9)	51673031.0(26)	R(34)	0.04989
1695.81601(8)	50839285.1(25)	P(34)	0.04768	1723.94354(9)	51682527.1(27)	R(35)	0.04791
1696.30203(8)	50853855.5(25)	P(33)	0.04950	1724.25778(9)	51691948.0(27)	R(36)	0.04587
1696.78562(8)	50868353.1(25)	P(32)	0.05124	1724.56952(9)	51701293.4(28)	R(37)	0.04381
1697.26677(8)	50882777.7(24)	P(31)	0.05289	1724.87873(9)	51710563.5(28)	R(38)	0.04172
1697.74549(8)	50897129.4(24)	P(30)	0.05442	1725.18543(10)	51719758.1(29)	R(39)	0.03963
1698.22178(8)	50911408.2(24)	P(29)	0.05583	1725.48960(10)	51728877.0(29)	R(40)	0.03755
1698.69564(8)	50925614.2(24)	P(28)	0.05709	1725.79126(10)	51737920.2(29)	R(41)	0.03549
1699.16707(8)	50939747.2(24)	P(27)	0.05820				continued

Table V, continued

	rable v, continue	<u> </u>	
			Intensity
***	**	TD.	$(m^{-2} daPa^{-1})$
Wave number	Frequency	Tran-	or
(cm <sup>-1</sup> )	(MHz)	sition	$({\rm cm}^{-2}{\rm atm}^{-1})^a$
1726.09038(10)	51746887.7(30)	R(42)	0.03345
1726.38697(10)	51755779.3(30)	R(43)	0.03145
1726.68103(10)	51764594.9(30)	R(44)	0.02950
1726.97255(10)	51773334.5(31)	R(45)	0.02760
1727.26153(10)	51781998.0(31)	R(46)	0.02576
1727.54797(10)	51790585.2(31)	R(47)	0.02399
1727.83186(10)	51799096.1(31)	R(48)	0.02228
1728.11321(10)	51807530.6(31)	R(49)	0.02065
1728.39200(10)	51815888.5(31)	R(50)	0.01909
1728.66823(11)	51824169.8(32)	R(51)	0.01761
1728.94191(11)	51832374.4(32)	R(52)	0.01620
1729.21302(11)	51840502.2(32)	R(53)	0.01487
1729.48157(11)	51848553.1(32)	R(54)	0.01362
1729.74755(11)	51856526.9(32)	R(55)	0.01245
1730.01095(11)	51864423.6(32)	R(56)	0.01135
1730.27178(11)	51872243.1(32)	R(57)	0.01032
1730.53003(11)	51879985.2(32)	R(58)	0.00937
1730.78570(11)	51887649.9(32)	R(59)	0.00848
1731.03878(11)	51895237.0(33)	R(60)	0.00766
1731.28926(11)	51902746.4(33)	R(61)	0.00691
1731.53716(11)	51910178.1(33)	R(62)	0.00622
1731.78245(11)	51917531.8(34)	R(63)	0.00558
1732.02515(11)	51924807.6(34)	R(64)	0.00500
1732.26523(12)	51932005.3(35)	R(65)	0.00446
1732.50271(12)	51939124.7(35)	R(66)	0.00398
1732.73758(12)	51946165.7(36)	R(67)	0.00354
1732.96983(12)	51953128.3(36)	R(68)	0.00314
1733.19945(12)	51960012.4(37)	R(69)	0.00278
1733.42645(12)	51966817.7(37)	R(70)	0.00246
1733.65082(12)	51973544.2(37)	R(71)	0.00217
1733.87256(12)	51980191.7(37)	R(72)	0.00191
1734.09166(12)	51986760.2(37)	R(73)	0.00168
1734.30812(12)	51993249.5(37)	R(74)	0.00147
1734.52194(12)	51999659.4(37)	R(75)	0.00129
1734.73310(12)	52005990.0(37)	R(76)	0.00112
1734.94161(12)	52012240.9(37)	R(77)	0.00098
1735.14746(12)	52018412.2(37)	R(78)	0.00085
1735.35065(12)	52024503.7(37)	R(79)	0.00074
1735.55117(12)	52030515.2(37)	R(80)	0.00064
1735.74902(13)	52036446.6(38)	R(81)	0.00055
1735.94420(13)	52042297.8(40)	R(82)	0.00047
1736.13669(14)	52048068.7(42)	R(83)	0.00041
1736.32651(15)	52053759.1(46)	R(84)	0.00035
1736.51363(17)	52059368.9(51)	R(85)	0.00030
1736.69806(19)	52064898.0(57)	R(86)	0.00026
1736.87979(22)	52070346.2(66)	R(87)	0.00022
1737.05882(25)	52075713.4(75)	R(88)	0.00019
1737.23515(29)	52080999.5(86)	R(89)	0.00016
1737.40876(33)	52086204.3(99)	R(90)	0.00013
1737.57966(38)	52091327.6(114)	R(91)	0.00011
1737.74783(43)	52096369.4(130)	R(92)	0.00010
1737.91328(49) 1738.07601(56)	52101329.5(148)	R(93)	0.00008
1738.23599(63)	52106207.8(168) 52111004.1(190)	$R(94) \\ R(95)$	0.00007
1100.20099(00)	52111004.1(190)	n (95)	0.00006

 $<sup>^</sup>a$  To within the accuracy of the intensity data, 1  $\rm m^{-2}\,daPa^{-1}$  is equal to 1  $\rm cm^{-2}\,atm^{-1}$  .

The line intensities for the  $2\nu_1$  band given in Table V were based on a value of 61 cm<sup>-2</sup> MPa<sup>-1</sup> (6.1 cm<sup>-2</sup> atm<sup>-1</sup>) for the integrated total band intensity as recently measured by Kagann.<sup>15</sup> This value is equivalent to a transition moment of 0.0183 D, which is close to 0.021 D measured by Foord and Whiffen.<sup>14</sup>

We have measured frequencies of the OCS spectrum at 11.6  $\mu$ m using laser frequency metrology techniques which permit better than an order of magnitude improvement in the constants used to calculate the frequencies. These predicted frequencies (for both the  $10^{0}0-00-00$  and  $20^{0}0-00^{0}0$  bands) are presented here for immediate use as reference standards. The ultimate goal of this and subsequent heterodyne measurements on the other bands is to facilitate an exact analysis of OCS as a whole in order to provide the best OCS standard frequencies possible.

We wish to express our deep appreciation to Andre Fayt of the University of Louvain, Louvain-la-Neuve, Belgium, for providing us some of his calculations on OCS line positions. We are also grateful to the NASA Upper Atmosphere Research Office for partial support of this program.

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# Heterodyne frequency measurements on the 11.6- $\mu$ m band of OCS: new frequency/wavelength calibration tables for 11.6-and 5.8- $\mu$ m OCS bands: erratum

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Received 26 May 1981.

The penultimate line in the body of Table IV of this paper<sup>1</sup> was garbled in the printing process. It is run below correctly.

Table IV. Constant a for OCS Determined from These Measurements

Constant	This work (MHz) <sup>b</sup>	Previous work <sup>4</sup> (MHz)
$\nu_0 (20^00-00^00)$	51 293 779.3 (19)	51 293 779. (111)
$\nu_0 (10^00-00^00)$	25 751 180.2 (6)	25 751 204. (65)
$v_0 (20^{0}0-10^{0}0)$	25 542 599.0 (18)	
$\nu_0 \ (11^10 - 01^10)$	25 543 432.7 (14)	25 543 403. (75)
$B(00^00)$	6081.492518 (56)	(6081.4924439)¢
D(0000)	$1.3019369 (114) \times 10^{-3}$	$(1.301789 \times 10^{-3})^c$
$B(10^00)$	6063.358265 (200)	6063.3588 (8)
$D(10^{0}0)$	$1.329873(197) \times 10^{-3}$	$1.3312(19) \times 10^{-3}$
$H(10^00)$	$0.0001214(276) \times 10^{-6}$	
$B(20^00)$	6044.874923 (2550)	6044.891 (5)
$D(20^{0}0)$	$1.359731 (810) \times 10^{-3}$	$1.3607 (10) \times 10^{-3}$
$H(20^{0}0)$	$0.0002234 (706) \times 10^{-6}$	-
$B(01^{1}0)$	6092.078222 (289)	_
$D(01^{1}0)$	$1.323567 (625) \times 10^{-3}$	
$q(01^10)$	6.361411394 (745)	
$q_{J}$ (01 <sup>1</sup> 0)	$4.25608 (667) \times 10^{-6}$	-
$B(11^{1}0)$	6075.50528 (186)	6075.519 (7)
$D(11^{1}0)$	$1.362242 (679) \times 10^{-3}$	
$q(11^10)$	6.853553 (2121)	6.853 (13)
$q_J (11^10)$	$11.670 (489) \times 10^{-6}$	

<sup>&</sup>lt;sup>a</sup> The statistically estimated uncertainties in the last digits (one standard deviation) are given in parentheses following each value. Since systematic errors have not been taken into account, these estimated uncertainties may be slightly optimistic.

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<sup>&</sup>lt;sup>b</sup> To convert from frequency units to wave number units, use 299 792 458 m/sec for the speed of light. <sup>16</sup>

 $<sup>^{</sup>c}$  See Ref. 10; the negligible difference in the  $B_{0}$  and  $D_{0}$  values given in Ref. 10 and in this work arises from the inclusion of an  $H_{0}$  term in Ref. 10.